Numerical assessment of convergence in the post-form Ichimura-Austern-Vincent model

Jin Lei 💿*

School of Physics Science and Engineering, Tongji University, Shanghai 200092, China

(Received 7 April 2025; accepted 13 May 2025; published 2 July 2025)

The Ichimura-Austern-Vincent (IAV) model provides a powerful theoretical framework for describing inclusive breakup reactions. However, its post-form representation presents significant numerical challenges due to the absence of a natural cutoff in the transition matrix integration. This work presents a systematic assessment of convergence methods for post-form IAV calculations, comparing the bin method and the Vincent-Fortune approach. I demonstrate that while the bin method offers implementation simplicity, it exhibits strong parameter dependence that compromises numerical stability. In contrast, the Vincent-Fortune method, which employs complex contour integration, achieves reliable convergence without arbitrary parameters. I further introduce a novel hybrid approach that integrates finite-range distorted wave Born approximation (DWBA) with the Vincent-Fortune technique, combining the accuracy of finite-range treatment at short distances with the numerical stability of zero-range approximations in the asymptotic region. Numerical results for deuteron and ⁶Li-induced reactions confirm the efficacy of this hybrid method, achieving reliable outcomes while resolving the convergence problems associated with the bin approach. This advancement enables more reliable calculations of nonelastic breakup cross sections and facilitates the extension of the IAV formalism beyond DWBA to incorporate continuum-discretized coupled-channels wave functions for a more comprehensive treatment of breakup processes.

DOI: 10.1103/153j-32cp

I. INTRODUCTION

In nuclear physics, the breakup process is a fundamental phenomenon wherein a projectile nucleus disintegrates into smaller fragments upon colliding with a target nucleus. This process is essential for probing nuclear structure, yielding critical insights into properties such as separation energies, angular momenta, parities, and electric responses to the continuum [1–6]. Beyond structural exploration, breakup studies illuminate the reaction mechanisms governing weakly bound systems, facilitate the calculation of incomplete fusion cross sections, enable the determination of neutron-induced cross sections for short-lived nuclei through surrogate reactions, and enhance our understanding of nuclear astrophysics and the synthesis of heavy isotopes [7–11].

For two-body projectiles, the breakup reaction is typically denoted as $a + A \rightarrow b + x + A$, where the projectile *a* comprises fragments *b* and *x*. When all three outgoing particles, *b*, *x*, and the target *A*, are detected in a well-defined final state, the reaction is classified as *exclusive*. Such reactions can be modeled as an effective three-body scattering problem using a suitable interaction potential. A rigorous theoretical approach to this problem is provided by the Faddeev equations [12–15], which can also incorporate excitations of the target or core. However, the computational complexity of solving these equations restricts their use to specific cases. As a result, more practical methods, such as the continuumdiscretized coupled-channels (CDCC) method [16,17] and the distorted wave Born approximation (DWBA) [18,19], have gained widespread adoption due to their balance of accuracy and feasibility.

In contrast, *inclusive* reactions occur when the final state is not fully specified. For example, in a reaction of the form A(a, bX), only one constituent of the projectile, say b, is observed, while X represents the unobserved particles, rendering the reaction inclusive with respect to those components. The simplest contribution to the inclusive cross section is elastic breakup (EBU), where all three outgoing particles remain in their ground states. However, more complex processes can also contribute, including breakup with excitation of x or A, particle transfer between x and A, or incomplete fusion, where x fuses with A. These additional processes collectively constitute nonelastic breakup (NEB).

Given the multitude of possible final states, explicitly computing the NEB contribution by summing over all processes is generally impractical. To address this, several research groups in the 1980s developed closed-form expressions that leverage the completeness of the x + A states to formally sum over final states [20–22]. Among these, the model proposed by Ichimura, Austern, and Vincent (IAV) has attracted significant interest in recent years. Multiple groups have revisited and implemented the IAV model, demonstrating promising agreement with experimental data, a distinction not consistently achieved by competing approaches [23–30].

The original IAV model employs closed-form expressions with post-form interactions. However, a notable limitation of this representation is the absence of a natural cutoff in the integration of the transition matrix, leading to nonconverging numerical results due to long-range effects. To mitigate this, Huby and Mines introduced a convergence factor, $e^{-\alpha r}$ [31], while Thompson proposed the bin method [32]. Both techniques, however, are highly sensitive to parameter choices, complicating efforts to achieve consistent convergence. Alternatively, Vincent and Fortune employed integration in the complex plane [33], though this method is limited to the zero-range approximation with finite-range corrections. Consequently, practical applications of the IAV model often rely on its equivalent prior form. While the prior form performs adequately within the DWBA framework, it proves insufficient for capturing higher-order effects, such as those in Austern's model using the three-body CDCC wave function, which I have recently implemented numerically with Moro [34,35], necessitating a return to the post-form.

In my previous work, we utilized the bin method to manage the post-form representation with both CDCC and DWBA wave functions [11,26,27,29,34,35]. However, the results were sensitive to the choice of matching radius and bin size, and convergence was not consistently achieved across all parameter values. To address these convergence challenges, this paper proposes a detailed comparison between the bin method and the Vincent-Fortune method. Furthermore, I introduce a novel approach that integrates the finite-range DWBA with the Vincent-Fortune method, aiming to enhance convergence and improve the reliability of practical calculations. This study seeks to refine the computational framework for breakup reactions, advancing both theoretical understanding and experimental validation in nuclear physics.

The paper is organized as follows. In Sec. II, I present the theoretical framework of the IAV model, focusing on the formulation of nonelastic breakup reactions and the challenges associated with the post-form representation. Section III introduces the convergence methods investigated in this work, including a detailed exposition of the novel hybrid approach that combines finite-range DWBA with the Vincent-Fortune technique. In Sec. IV, I present numerical applications to deuteron and ⁶Li-induced reactions, systematically comparing the performance of the bin method and the Vincent-Fortune method. Section V offers a critical discussion of the strengths and limitations of each method, with particular emphasis on their applicability to more complex systems involving CDCC wave functions. Finally, Sec. VI summarizes the findings and outlines directions for future research.

II. THEORETICAL FRAMEWORK

The process under study involves a projectile labeled a, which has a two-body structure denoted as a = b + x. This projectile collides with a target nucleus A, leading to the emission of a fragment b. In this reaction, b acts as a spectator, while x is the participant interacting with A. The process can be represented as

$$a(=b+x)+A \to b+B^*,$$

where B^* denotes any resulting state of the x + A system.

Applying energy conservation in Jacobi coordinates, one obtains

$$E_{bx} + E_a = E_b + E_x,$$

where E_{bx} is the relative energy of the b + x pair, E_x is the relative energy of the x + A pair, E_a is the relative energy of a with respect to A, and E_b is the relative energy of b with respect to the x + A system. The corresponding wave numbers for E_a , E_b , and E_x are denoted k_a , k_b , and k_x , respectively.

The interaction between x and A encompasses both elastic scattering and nonelastic reactions. The elastic scattering is termed EBU, while the nonelastic processes, collectively referred to as NEB, include inelastic scattering of x + A, nucleon exchange between x and A, fusion, and transfer to bound states of B.

In the three-body model proposed by IAV, the differential cross section for the NEB inclusive process is given by the closed-form expression

$$\frac{d^2\sigma}{dE_b d\Omega_b} \bigg|_{\text{NEB}} = -\frac{2}{\hbar v_a} \rho_b(E_b) \langle \varphi_x(\mathbf{k}_b) | \text{Im}[U_{xA}] | \varphi_x(\mathbf{k}_b) \rangle, \quad (1)$$

where $\rho_b(E_b)$ is the density of states of particle *b*, v_a is the velocity of the incoming projectile *a*, $\varphi_x(\mathbf{k}_b, \mathbf{r}_{xA})$ is the relative wave function describing the motion between *x* and *A* when *b* is scattered with momentum \mathbf{k}_b , and U_{xA} is the effective optical potential between *x* and *A*. The wave function $\varphi_x(\mathbf{k}_b, \mathbf{r}_x)$ is determined by

$$\varphi_x(\mathbf{k}_b, \mathbf{r}_x) = \int_0^{R_{\text{max}}} G_x(\mathbf{r}_x, \mathbf{r}'_x) \langle \mathbf{r}'_x | \rho \rangle \, d\mathbf{r}'_x, \qquad (2)$$

where the source term is

$$\langle \mathbf{r}_{x} | \rho \rangle = \langle \mathbf{r}_{x} \chi_{b}^{(-)}(\mathbf{k}_{b}) | \mathcal{V}_{\text{post}} | \Psi^{3b(+)} \rangle.$$
(3)

Here, G_x is the Green's function incorporating the optical potential U_{xA} , $\chi_b^{(-)*}(\mathbf{k}_b, \mathbf{r}_b)$ is the distorted wave describing the relative motion of *b* with respect to the B^* system (obtained using an optical potential U_{bB}), and $\Psi^{3b(+)}$ is the exact three-body scattering wave function. The post-form transition operator is defined as $\mathcal{V}_{\text{post}} = V_{bx} + U_{bA} - U_{bB}$, where V_{bx} is the binding potential of the b + x projectile, and U_{bA} is the optical potential for the relative scattering of *b* and *A*. In the asymptotic limit, the wave function $\varphi_x(\mathbf{k}_b, \mathbf{r}_x)$ takes the form

$$\varphi_x(\mathbf{k}_b, \mathbf{r}_x) \underset{r_x \to \infty}{\longrightarrow} f(\hat{k}_b, \hat{r}_x) \frac{e^{ik_x r_x}}{r_x}, \tag{4}$$

where $f(\hat{k}_b, \hat{r}_x)$ is the scattering amplitude, which can be used to compute the EBU cross section, as detailed in Ref. [26].

In the three-body model, the exact three-body wave function $\Psi^{3b(+)}$ is often approximated using the DWBA, expressed as $\Psi^{\text{DWBA}(+)} = \chi_a^{(+)}\phi_a$, where $\chi_a^{(+)}$ describes the elastic scattering of a + A, and ϕ_a is the bound-state wave function of the projectile *a*. Although the original IAV model employed the DWBA framework [20], Austern *et al.*[17] extended it by incorporating the CDCC method, which includes continuum states of the b + x system. For simplicity, this paper focuses on the DWBA form, which has been shown to be a reliable approximation of the CDCC wave function [11,34,35]. In configuration space, evaluating Eq. (2) poses numerical challenges, often leading to divergence. This issue arises because V_{bx} , the dominant component of \mathcal{V}_{post} , shares the same spatial coordinate as ϕ_a , resulting in an unbounded r_x integration without a natural cutoff. To address this, Huby and Mines [31] and Vincent [36] introduced a convergence factor, redefining the source term as

$$\langle \mathbf{r}_{x} | \rho \rangle \equiv \lim_{\alpha \to 0^{+}} \langle \mathbf{r}_{x} | e^{-\alpha \mathbf{r}_{x}} \rho \rangle.$$
 (5)

The damping factor $e^{-\alpha \mathbf{r}_x}$ ensures numerical stability. In principle, converged results are obtained by computing $\langle \mathbf{r}_x | e^{-\alpha \mathbf{r}_x} \rho \rangle$ for various α values and extrapolating to the limit $\alpha \to 0^+$. In practice, however, a small fixed α is typically chosen, making the numerical results sensitive to this parameter.

Another possible solution to this convergence problem is the bin method proposed by Thompson [32], which rewrites the source term as

$$\langle \mathbf{r}_{x} | \rho \rangle \approx \langle \chi_{b}^{(-)}(\mathbf{k}_{b}) \big| \chi_{b}^{\text{bin}}(\bar{k}_{b}) \big\rangle \langle \mathbf{r}_{x} \chi_{b}^{\text{bin}}(\bar{k}_{b}) \big| \mathcal{V}_{\text{post}} | \Psi^{3b(+)} \rangle, \quad (6)$$

with

$$\left\langle \chi_{b}^{\mathrm{bin}}(\bar{k}_{b}) \middle| \mathbf{r}_{b} \right\rangle = \sqrt{\frac{2}{\pi (k_{n+1} - k_{n})}} \int_{k_{n}}^{k_{n+1}} \langle \chi_{b}^{(-)}(\mathbf{k}_{b}) | \mathbf{r}_{b} \rangle \, d\mathbf{k}_{b}.$$
(7)

Since the bin wave function $\chi_b^{\text{bin}}(\bar{k}_b)$ is square integrable, it resolves the convergence problem in Eq. (2). However, the starting point of Eq. (6) involves inserting a complete basis of bin states and assuming this basis approximates the unit operator. With energy conservation, only the bin state covering the scattering state $\chi_b^{(-)}(\mathbf{k}_b)$ remains. Thus, the numerical results strongly depend on the bin size $\Delta k = k_{n+1} - k_n$. When the bin size is too small, a large integration interval is required to achieve convergence. Conversely, if the bin size is too large, the assumption that the basis of bin states equals the unit operator may not hold, potentially leading to misleading results. Therefore, achieving convergence requires extensive parameter tuning, balancing the matching radius and bin size.

Vincent and Fortune [33] introduce a distinct approach to transform the problematic real-axis integral into a more manageable form by employing contour integration in the complex radius plane. To achieve this, they divide the integration range into two regions: an inner region, where the integration can be evaluated numerically using standard methods, and an outer region, where the results are obtained through complex contour integration. Unlike the scaling factor method or the bin method, the Vincent-Fortune method is parameter free, meaning its numerical results should not depend on arbitrarily chosen parameters. However, it requires special treatment of the outer region, where the scattering wave function is expressed in terms of outgoing (H^+) and incoming (H^-) Hankel functions. Consequently, the Vincent-Fortune method is currently limited to the zero-range approximation with finite-range corrections. Extending it to the finite-range form necessitates a coordinate transformation from incoming to outgoing Jacobi coordinates, which significantly complicates the handling of Hankel functions across different Jacobi coordinate systems.

III. CONVERGENCE METHOD

In this paper, I present a novel method for computing the relative wave function $\varphi_x(\mathbf{k}_b, \mathbf{r}_x)$ in Eq. (2), building upon the Vincent-Fortune approach. The reaction that is being studied here occurs in a regime where short-range nuclear forces dominate, typically acting over distances of just a few femtometers. These forces bind the projectile a = b + x, and the zero-range approximation simplifies this interaction by treating the potential V_{bx} as a delta function, meaning b and x only interact when they are at the same point in space. While this assumption is a simplification, it becomes increasingly reasonable at large radial distances $(r_x \to \infty)$, where the finite-range DWBA source term $\langle \mathbf{r}_x | \rho \rangle$, which captures the extended spatial structure of the b + x system, approaches the simpler zero-range form. This behavior aligns with the asymptotic limit of the wave function, where $\varphi_x(\mathbf{k}_b, \mathbf{r}_x) \propto e^{ik_x r_x}/r_x$, reflecting the outgoing scattering of x relative to A.

Inspired by this observation, the proposed method splits the computation of $\varphi_x(\mathbf{k}_b, \mathbf{r}_x)$ into two regions, leveraging the strengths of both finite-range and zero-range descriptions. Specifically, I express the wave function as

$$\varphi_{x}(\mathbf{k}_{b},\mathbf{r}_{x}) \approx \int_{0}^{R} G_{x}(\mathbf{r}_{x},\mathbf{r}_{x}') \langle \mathbf{r}_{x}' | \rho^{\text{exact}} \rangle \, d\mathbf{r}_{x}' + \int_{R}^{\infty} G_{x}(\mathbf{r}_{x},\mathbf{r}_{x}') \langle \mathbf{r}_{x}' | \rho^{\text{zero}} \rangle \, d\mathbf{r}_{x}', \qquad (8)$$

where $G_x(\mathbf{r}_x, \mathbf{r}'_x)$ is the Green's function incorporating the optical potential U_{xA} , $\langle \mathbf{r}'_x | \rho^{\text{exact}} \rangle$ is the full finite-range DWBA source term, and $\langle \mathbf{r}'_x | \rho^{\text{zero}} \rangle$ is its zero-range counterpart. The boundary *R* separates the inner region (from zero to *R*), where nuclear interactions are strong and the finite-range term is essential, from the outer region (from *R* to ∞), where the zero-range approximation suffices due to the diminishing influence of V_{bx} . In the outer region, the integration is performed using the complex contour techniques of Vincent and Fortune, capitalizing on the asymptotic simplicity of the zero-range source term. The detailed partial wave decomposition of the source term can be found in Ref. [37], and the partial wave decomposition of the zero-range source and the implementation of the Vincent-Fortune method can be found in the Appendix.

This hybrid approach offers clear advantages. The use of the finite-range DWBA source term up to R ensures accuracy in the inner region, where nuclear interactions shape the x + A dynamics at short distances. Meanwhile, the zero-range source term enables efficient complex contour integration in the outer region, where long-range scattering dominates, avoiding the divergence issues of unbounded realaxis integrals. By integrating the parameter-free framework of the Vincent-Fortune method with a practical handling of finite-range effects, this new approach provides a robust and physically sound solution to the convergence challenges in Eq. (2). It ensures numerical stability without relying on *ad* hoc parameters like damping factors or bin sizes, ultimately enabling more precise calculations of breakup cross sections, such as those for elastic and nonelastic processes. This method thus represents a significant step forward in modeling threebody nuclear reactions with greater accuracy and reliability.



FIG. 1. Radial parts of the bin state wave function at a proton energy of 14 MeV in the laboratory frame, calculated with different bin sizes: (a) $\Delta k = 0.08 \text{ fm}^{-1}$, (b) $\Delta k = 0.06 \text{ fm}^{-1}$, (c) $\Delta k = 0.04 \text{ fm}^{-1}$, and (d) $\Delta k = 0.02 \text{ fm}^{-1}$.

IV. NUMERICAL RESULTS

In this section, I present calculations for reactions induced by deuteron and ⁶Li projectiles and compare the calculated NEB cross section with the post-form IAV model using the bin method and the Vincent-Fortune method. In all cases, I ignore the spin of the particles to simplify the calculations.

A. Application to (d, pX)

As a first example, I consider the reaction ${}^{62}\text{Ni}(d, pX)$ at an incident deuteron energy $E_d = 25.5$ MeV, which was previously analyzed in Ref. [27].

In the calculations, the deuteron ground-state wave function was generated using a simple Gaussian potential, as described in Ref. [17]. The deuteron distorted waves were computed with an optical potential sourced from Ref. [38]. For the proton-target and neutron-target interactions, I employed the global parametrization of Koning and Delaroche from Ref. [39].

As outlined in the previous section, evaluating the postform formulas requires specialized techniques. Here, I first discuss the bin method, which involves averaging the distorted waves $\chi_b^{(-)}(\mathbf{k}_b, \mathbf{r}_b)$ over small momentum intervals (bins). This averaging renders the resulting functions square integrable, making the source term in Eq. (6) short ranged. Before presenting the NEB results from the post-form IAV model, I examine the radial part of the bin state wave function at a proton energy of 14 MeV in the laboratory frame, as shown in Fig. 1. The figure displays the radial components for partial waves with $0 \le \lambda_b \le 25$, calculated using bin sizes of $\Delta k = 0.08$, 0.06, 0.04, and 0.02 fm⁻¹, labeled as (a), (b), (c), and (d), respectively.

The figure reveals that as the bin size (Δk) increases, the bin state wave function becomes more spatially compressed, localizing within a smaller region. For example, at $\Delta k = 0.08 \text{ fm}^{-1}$, individual wave packets are distinguishable, with



FIG. 2. Differential cross section energy distribution of the partial wave components of the NEB for the reaction $^{62}Ni(d, pX)$ at $E_d = 25.5$ MeV, with the outgoing proton energy at 14 MeV. (a) Fixed $\Delta k = 0.08$ fm⁻¹ with varying R_{max} values, where R_{max} is the maximum radial distance in the numerical calculations. (b) Results with different Δk values, each with R_{max} set at the nodes of the outgoing proton wave functions. See text for details.

the amplitudes of subsequent packets decreasing gradually until a square-integrable wave function emerges. This behavior aligns with the uncertainty principle: a larger Δk corresponds to a smaller uncertainty in position, Δr . Conversely, at $\Delta k =$ 0.02 fm^{-1} , the wave packet exhibits significant spatial broadening, resulting in a larger spread in position space.

In principle, achieving fully converged results with the bin method requires setting the maximum integration range, R_{max} , to a value where the radial part of the bin state wave function effectively vanishes. However, this ideal scenario presents significant numerical challenges. For instance, as shown in Fig. 1, even with the largest bin size considered ($\Delta k = 0.08 \text{ fm}^{-1}$), an R_{max} of 300 fm is insufficient to ensure the wave function fully diminishes.

To investigate this further, I analyze the differential cross section energy distribution for the partial wave components of the NEB in the reaction 62 Ni(*d*, *pX*) at $E_d = 25.5$ MeV, with the outgoing proton energy fixed at 14 MeV in the center-of-mass frame. These results are presented in Fig. 2(a), where calculations are performed with a fixed bin size of $\Delta k = 0.08$ fm⁻¹ and varying R_{max} values: 80 fm (solid line),



FIG. 3. Differential cross section energy distribution of the NEB for the reaction ${}^{62}\text{Ni}(d, pX)$ at $E_d = 25.5$ MeV as a function of the outgoing proton energy. Results are obtained using different Δk values, with R_{max} set at the nodes of the outgoing proton wave functions for each Δk . See text for details.

100 fm (dotted line), 120 fm (dashed line), and 155 fm (dash-dotted line). For the partial wave $\lambda_a = 10$, significant discrepancies appear across these R_{max} values, indicating that simply increasing R_{max} does not guarantee convergence.

In practice, the bin method often sets R_{max} at the nodes of the bin state wave function to confine it within a wave packet. For $\Delta k = 0.08 \text{ fm}^{-1}$, these nodes occur at approximately 80, 155, and 235 fm, and beyond. However, the results show a marked difference between $R_{\text{max}} = 80$ and 155 fm, highlighting numerical instability in the bin method. This suggests that relying solely on node-based R_{max} selections does not consistently yield stable, converged results.

Figure 2(b) presents results for different Δk values, with R_{max} set to the first node of the outgoing proton wave function for each case: $\Delta k = 0.08 \text{ fm}^{-1}$ with $R_{\text{max}} = 80 \text{ fm}$ (solid line), $\Delta k = 0.06 \text{ fm}^{-1}$ with $R_{\text{max}} = 105 \text{ fm}$ (dotted line), $\Delta k = 0.04 \text{ fm}^{-1}$ with $R_{\text{max}} = 155 \text{ fm}$ (dashed line), and $\Delta k = 0.02 \text{ fm}^{-1}$ with $R_{\text{max}} = 310 \text{ fm}$ (dash-dotted line). Noticeable differences are evident among the results for $\Delta k = 0.08, 0.06$, and 0.04 fm⁻¹. Although larger Δk values compress the wave function into a smaller spatial region, these bin sizes appear too coarse to accurately capture the features of the original scattering wave function. However, the results for $\Delta k = 0.04 \text{ and } 0.02 \text{ fm}^{-1}$ show negligible differences, suggesting convergence, consistent with findings in prior studies (e.g., Refs. [11,26,27,29,34,35]).

Next, I present the differential cross section energy distribution of the NEB as a function of the outgoing proton energy in the center-of-mass frame, shown in Fig. 3. The results correspond to $\Delta k = 0.08 \text{ fm}^{-1}$ with $R_{\text{max}} = 80$ fm (solid line), $\Delta k = 0.06 \text{ fm}^{-1}$ with $R_{\text{max}} = 105$ fm (dotted line), $\Delta k = 0.04 \text{ fm}^{-1}$ with $R_{\text{max}} = 155$ fm (dashed line), and $\Delta k = 0.02 \text{ fm}^{-1}$ with $R_{\text{max}} = 310$ fm (dash-dotted line). A vertical line marks the threshold where the relative energy between the neutron and ⁶²Ni is zero. The differences between these calculations are more pronounced than those in





FIG. 4. Absolute value of the difference of the Λ function [see Eq. (A16)], computed using finite-range and zero-range DWBA for the reaction ⁶²Ni(*d*, *pX*) at $E_d = 25.5$ MeV, with an outgoing proton energy of 14 MeV in the center-of-mass frame, for partial waves with $l_x = 0, 0 \leq \lambda_a \leq 25$, and $0 \leq \lambda_b \leq 25$.

Fig. 2(b). This arises because the positions of the first nodes of the scattering wave function in coordinate space shift as the outgoing proton energy changes. While the calculations in Fig. 3 use a fixed R_{max} for simplicity, this approach proves inadequate for certain Δk values in this context.

The choice of R_{max} in the bin method remains ambiguous and even large values like 300 fm fail to ensure wave function decay. Convergence, observed only at $R_{\text{max}} = 155$ and 310 fm for smaller Δk , highlights the practical need for very large R_{max} values, posing computational challenges and underscoring the method's sensitivity to this parameter.

I now discuss the Vincent-Fortune method, which integrates finite-range and zero-range treatments of the source term function. The key assumption is that beyond a specific radial distance R, the finite-range and zero-range approximations produce equivalent results. For the present calculations, I adopt R = 80 fm. The source term function in the partial wave basis can be expressed as a product of radial and angular components. The angular components remain identical for both zero-range and finite-range treatments, whereas the radial components differ. I denote the radial part of the source term function by Λ [for the zero-range approximation, see Eq. (A16) for details]. To evaluate the validity of the zerorange approximation, Fig. 4 illustrates the absolute difference between the Λ function computed using the exact finite-range and zero-range approximations for partial waves with $l_x =$ 0, $0 \leq \lambda_a \leq 25$, and $0 \leq \lambda_b \leq 25$. The difference decreases exponentially with increasing radial distance, confirming the suitability of the zero-range approximation in Eq. (8).

One should also note that, for most zero-range calculations, a finite-range correction factor is typically defined, such as the one discussed in the Appendix of Ref. [25]. However, this correction factor only plays an important role within the nuclear interaction region. At large distances, it contributes only a constant factor. As shown in Fig. 4, without this factor, the difference between the zero-range and finite-range approximations decreases exponentially, confirming the validity



FIG. 5. Integrand function, fixed at x = R on the complex plane, varying along the y axis for the reaction ${}^{62}\text{Ni}(d, pX)$ at $E_d = 25.5$ MeV, with an outgoing proton energy of 14 MeV, for partial waves with $l_x = \lambda_a = \lambda_b = 0$. See text for details.

of the zero-range approximation at large distances. For this reason, I did not include the finite-range correction factor in the current calculations.

To assess the numerical stability of the Vincent-Fortune method, I plot the integrand functions from Eq. (A27) in Fig. 5 for the same reaction, using partial waves with $l_x = \lambda_a = \lambda_b = 0$. The figure comprises two panels. In the upper panel, the solid line represents $U1 = |H_{l_x}^{(+)}(k_xz)H_{\lambda_b}^{(-)}(k_bcz)H_{\lambda_a}^{(+)}(k_az)|$, and the dotted line shows $U2 = |H_{l_x}^{(+)}(k_xz)H_{\lambda_b}^{(+)}(k_bcz)H_{\lambda_a}^{(+)}(k_az)|$. In the lower panel, the dashed line depicts $L1 = |H_{l_x}^{(+)}(k_xz)H_{\lambda_b}^{(-)}(k_bcz)H_{\lambda_a}^{(-)}(k_az)|$, and the dash-dotted line represents $L2 = |H_{l_x}^{(+)}(k_xz)H_{\lambda_b}^{(-)}(k_bcz)H_{\lambda_a}^{(-)}(k_az)|$. All four components (U1, U2, L1, and L2) are smooth functions. Three of them (U1, U2, and L1) are short ranged, with significant contributions only up to approximately 7 fm. In contrast, L2 is long ranged, extending along the y direction to about 70 fm. This behavior stems from the exponential growth of $H_{l_x}^{(+)}(k_xz)$ and $H_{\lambda_b}^{(+)}(k_bcz)$ in the lower panel, partially offset by the faster exponential decay of $H_{\lambda_a}^{(-)}(k_az)$. However, this decay does not fully compensate for the growth, slowing the convergence of L2 and allowing it to extend over a larger distance.

Finally, I compare the results from the bin method and the Vincent-Fortune method for the same reaction in Fig. 6. The solid line represents the bin method with $\Delta k = 0.02 \text{ fm}^{-1}$, while the dashed line corresponds to the Vincent-Fortune method. The two sets of results show reasonable agreement. However, unlike the bin method, which depends on parameters like Δk and R_{max} , the Vincent-Fortune method is parameter independent and exhibits better numerical stability and convergence.

B. Application to $(^{6}Li, \alpha X)$

As a second example, I examine the reaction ²⁰⁹Bi(⁶Li, αX), which was previously analyzed in Ref. [26] using the



FIG. 6. Differential cross section energy distribution of the NEB for the reaction ${}^{62}\text{Ni}(d, pX)$ at $E_d = 25.5$ MeV as a function of the outgoing proton energy. Results are obtained using the bin method and the Vincet-Fortune method.

post-form IAV model with the bin method. Those calculations successfully reproduced the experimental angular distribution of α particles across a wide range of incident energies, both above and below the Coulomb barrier. To assess the numerical stability of the bin method and compare its results with those obtained using the Vincent-Fortune approach, I consider an incident energy of E = 50 MeV. For the calculations presented here, I adopt the potentials used in Ref. [26].

I first compare the differential cross-section energy distributions computed using the bin method with varying Δk values, as shown in Fig. 7(a), for the reaction ²⁰⁹Bi(⁶Li, αX) at $E_{lab} = 50$ MeV. The solid, dotted, dashed, and dash-dotted lines represent results with $\Delta k = 0.08$ fm⁻¹ and $R_{max} = 80$ fm, $\Delta k = 0.06$ fm⁻¹ and $R_{max} = 110$ fm, $\Delta k = 0.04$ fm⁻¹ and $R_{max} = 155$ fm, and $\Delta k = 0.02$ fm⁻¹ and $R_{max} = 310$ fm, respectively. Unlike the deuteron case, a clear trend toward convergence emerges as Δk decreases, suggesting that the bin method is numerically unreliable. In Fig. 7(b), I present a comparison for the same reaction using the Vincent-Fortune method. The solid and dashed lines correspond to results from the bin method with $\Delta k = 0.02$ fm⁻¹ and $R_{max} = 310$ fm and the Vincent-Fortune method, respectively. In contrast to the bin method, the Vincent-Fortune method exhibits significantly greater numerical stability.

V. DISCUSSION

As discussed above, the bin method, which averages distorted waves over momentum bins, is simple, flexible, and straightforward to implement by replacing the standard scattering wave function with a bin-state wave function. However, it is numerically unstable, exhibiting sensitivity to parameters such as R_{max} (the maximum radial cutoff) and Δk (the momentum bin size). This sensitivity leads to inconsistent results and poor convergence unless impractically small bin sizes and large radial ranges are employed. In contrast, the Vincent-Fortune method, which combines finite- and zero-



FIG. 7. Differential cross section energy distribution of the NEB for the reaction 209 Bi(⁶Li, αX) at $E_{lab} = 50$ MeV as a function of the outgoing proton energy. Results are obtained using the bin method and the Vincent-Fortune method.

range treatments, ensures rapid convergence and robustness without relying on arbitrary parameters. However, it assumes that the source term remains consistent beyond a radial distance R, an assumption that may not hold universally and is strictly valid only for *s*-wave bound-state projectiles. Additionally, its reliance on complex-plane integration can be computationally intensive. In practice, however, calculations using the prior-form IAV model can readily achieve convergence without numerical difficulties.

Challenges arise when extending scattering models beyond the DWBA, such as when incorporating CDCC wave functions into the IAV model, as explored in Refs. [11,34,35]. In such cases, ensuring convergence within the post-form formalism becomes critical for accurate predictions.

The CDCC wave function is typically expressed as a sum over discrete channels:

$$\Psi^{\text{CDCC}(+)}(\mathbf{r}_{bx},\mathbf{r}_{a}) = \sum_{i} \chi_{i}(\mathbf{r}_{bx})\phi_{i}(\mathbf{r}_{a}), \qquad (9)$$

where \mathbf{r}_{bx} denotes the relative coordinate between the projectile's constituents (e.g., a core and valence particle), \mathbf{r}_a is the coordinate between the projectile's center of mass and the target, $\chi_i(\mathbf{r}_{bx})$ describes the relative motion of the projectile-

target system in channel *i*, and $\phi_i(\mathbf{r}_a)$ represents the internal state of the projectile. For the ground state (*i* = 0), the term $\chi_0(\mathbf{r}_{bx})\phi_0(\mathbf{r}_a)$ corresponds to the DWBA, where χ_0 describes elastic scattering between the projectile and the target.

At large separations $(\mathbf{r}_a \to \infty)$, where the projectile and target are well separated, the three-body boundary condition governs the system's behavior. Under this condition, the wave function should describe elastic scattering, with the projectile remaining in its ground state (i = 0). Consequently, contributions from excited or continuum states (i > 0) diminish, and the CDCC wave function simplifies to predominantly the ground-state component.

Separately, the zero-range approximation can be applied to CDCC wave functions for two key reasons. First, reaction systems requiring CDCC typically involve weakly bound projectiles with *s*-wave ground states, which supports the use of a zero-range source term due to the simplicity of the groundstate wave function. Second, at large \mathbf{r}_a , the CDCC wave function approaches the elastic scattering form

$$\Psi^{\text{CDCC}(+)}(\mathbf{r}_a) \xrightarrow{\mathbf{r}_a \to \infty} \chi_0(\mathbf{r}_a).$$
(10)

Additionally, the relative coordinate \mathbf{r}_{bx} is effectively truncated because the interaction V_{bx} (between the projectile's constituents) is short ranged. As a result, the source term computed using the CDCC wave function converges toward the zero-range approximation at large distances. Mathematically, this implies

$$\rho^{\text{CDCC}} = \rho^{\text{DWBA}} \approx \rho^{\text{zero}}.$$
 (11)

Future work should focus on testing the Vincent-Fortune method in complex systems (e.g., exotic nuclei or higherenergy reactions), optimizing its computational efficiency, and integrating it with CDCC wave functions. The Vincent-Fortune method represents a reliable, parameter-independent improvement over the bin method, offering significant potential for precision studies of breakup reactions when combined with CDCC wave functions.

VI. SUMMARY AND CONCLUSIONS

In this paper, I have presented a comprehensive investigation of numerical convergence methods for the post-form IAV model, focusing on the calculation of nonelastic breakup reactions. The primary challenge addressed is the numerical instability inherent in the post-form representation, which lacks a natural cutoff in the integration of the transition matrix. Two main approaches were compared: the bin method and the Vincent-Fortune method, along with a novel hybrid approach that combines the finite-range DWBA with the Vincent-Fortune technique.

The bin method, which averages distorted waves over momentum bins to render them square integrable, offers a straightforward implementation but exhibits significant parameter sensitivity. As demonstrated in calculations for the 62 Ni(*d*, *pX*) reaction at $E_d = 25.5$ MeV, the results depend critically on both the bin size (Δk) and the maximum integration radius (R_{max}). While convergence can be achieved with sufficiently small bin sizes (around $\Delta k = 0.02$ fm⁻¹), this requires impractically large integration radii (over 300 fm), introducing computational inefficiencies. The method's reliance on setting R_{max} at the nodes of the bin state wave function further complicates its application, as these nodes shift with changing outgoing particle energies.

In contrast, the Vincent-Fortune method, which employs complex contour integration, demonstrates superior numerical stability without dependence on arbitrary parameters. The key insight underlying this approach is that beyond a certain radial distance, the finite-range and zero-range treatments of the source term yield equivalent results, as confirmed by the exponential decrease in their difference with increasing radius. The method divides the calculation into an inner region, where finite-range effects are essential, and an outer region, where zero-range approximations suffice, leveraging the asymptotic behavior of the wave function.

The proposed hybrid approach combines the strengths of both finite-range and zero-range descriptions, offering a robust solution to the convergence challenges in the post-form IAV model. For the reactions studied, ⁶²Ni(*d*, *pX*) at $E_d = 25.5$ MeV and ²⁰⁹Bi(⁶Li, αX) at E = 50 MeV, this approach produces results consistent with the bin method at small bin sizes, but with greater numerical stability and computational efficiency.

Despite these advances, several challenges remain. The Vincent-Fortune method assumes that the source term remains consistent beyond a certain radial distance, an assumption that may not hold universally and is strictly valid only for *s*-wave bound-state projectiles. Additionally, complex-plane integration can be computationally intensive, though this is offset by the method's rapid convergence. For practical applications, the prior-form IAV model often remains a viable alternative, as it naturally achieves convergence without numerical difficulties.

Looking forward, the convergence techniques developed here hold particular significance for extending the IAV model beyond the DWBA framework. The integration of the Vincent-Fortune method with CDCC wave functions, as explored in recent work [11,34,35], represents a promising direction for future research. This is especially relevant for systems involving weakly bound projectiles with *s*-wave ground states, where the zero-range approximation can be justified even for CDCC wave functions.

In conclusion, the hybrid approach combining finite-range DWBA with the Vincent-Fortune method represents a significant advancement in modeling three-body nuclear reactions. By providing a parameter-independent, physically motivated solution to the convergence challenges in the post-form IAV model, this method enhances the precision and reliability of breakup cross-section calculations, with broad implications for nuclear reaction studies.

ACKNOWLEDGMENTS

I am grateful to Antonio Moro for his invaluable discussions. This work has been partially supported by the National Natural Science Foundation of China (Grants No. 12475132 and No. 12105204) and by the Fundamental Research Funds for the Central Universities.

DATA AVAILABILITY

No data were created or analyzed in this study.

APPENDIX: PARTIAL WAVE ANALYSIS OF THE ZERO-RANGE SOURCE TERM AND VINCENT-FORTUNE METHOD

Instead of employing a three-dimensional Jacobi basis, the wave function can be expanded into partial wave eigenstates, which depend on the radial magnitude and angular momentum eigenstates. The orbital angular momenta of the three particles are coupled to a total angular momentum J and its third component M_J . For the incoming channels, the state is expressed as

$$|r_{bx}r_a\alpha_{\rm in}M_J\rangle = |r_{bx}r_a((l_a(j_bj_x)s_{bx})J_a(\lambda_aj_A)J_A)JM_J\rangle, \quad (A1)$$

and for the outgoing channels, it is given by

$$|r_x r_b \alpha_{\text{out}} M_J \rangle = |r_x r_b((l_x (j_x j_A) s_{xA}) J_x (\lambda_b j_b) J_b) J M_J \rangle, \quad (A2)$$

where j_b , j_x , and j_A denote the internal spins of particles b, x, and A, respectively; s_{bx} and s_{xA} represent the total spins of the *b*-*x* and *x*-*A* subsystems in the incoming and outgoing channels, respectively; l_a , λ_a , l_x , and λ_b are the relative orbital angular momenta of the *b*-*x*, *a*-*A*, *x*-*A*, and *b*-*B*^{*} pairs, respectively; and J_a (J_A) and J_x (J_b) are the total angular momenta of the subsystem (and spectator) in the incoming and outgoing channels, respectively.

The angular momentum basis can be normalized as follows:

$$\langle r'_{bx}r'_{a}\alpha'_{\rm in}|r_{bx}r_{a}\alpha_{\rm in}\rangle = \frac{\delta(r'_{bx} - r_{bx})}{r'_{bx}r_{bx}}\frac{\delta(r'_{a} - r_{a})}{r'_{a}r_{a}}\delta_{\alpha'_{\rm in},\alpha_{\rm in}}, \quad (A3)$$

with a similar normalization applying to the outgoing basis. Additionally, a two-body angular momentum basis for the x-A subsystem is defined as

$$|r_x\beta M_x\rangle = |r_x(l_x s_{xA})J_x M_x\rangle. \tag{A4}$$

Consequently, the three-body outgoing state can be decoupled into a product of subsystem states:

$$|r_{x}r_{b}\alpha_{\text{out}}M_{J}\rangle = \sum_{M_{x}M_{b}} \langle J_{x}M_{x}J_{b}M_{b}|JM_{J}\rangle |r_{x}\beta M_{x}\rangle |r_{b}J_{b}M_{b}\rangle,$$
(A5)

and similarly, the incoming state can be written as

$$|r_{bx}r_{a}\alpha_{\rm in}M_{J}\rangle = \sum_{M_{a}M_{A}} \langle J_{a}M_{a}J_{A}M_{A}|JM_{J}\rangle |r_{bx}J_{a}M_{a}\rangle |r_{a}J_{A}M_{A}\rangle,$$
(A6)

where M_x , M_b , M_a , and M_A are the third components of J_x , J_b , J_a , and J_A , respectively.

The source term in the IAV model, within the zero-range approximation in the partial-wave basis, can be expressed as

$$\langle r_{x}\beta M_{x}m_{j_{b}}|\rho M_{a}m_{j_{A}}\rangle = \langle r_{x}\beta M_{x}m_{j_{b}}\chi_{b}^{(-)}|V_{bx}|\chi_{a}^{(+)}\phi_{a}M_{a}m_{j_{A}}\rangle$$

$$= \sum_{\alpha_{out},\alpha_{in}}\sum_{M_{J}}\sum_{M_{b}}\sum_{M_{A}}\int r_{b}^{2}r_{a}^{2}r_{bx}^{2}dr_{a}dr_{b}dr_{bx}\langle J_{x}M_{x}J_{b}M_{b}|JM_{J}\rangle\langle J_{a}M_{a}J_{A}M_{A}|JM_{J}\rangle\delta_{\beta,\alpha_{out}}$$

$$\times \langle \chi_{b}^{(-)}m_{j_{b}}|r_{b}J_{b}M_{b}\rangle \int \mathcal{G}_{\alpha_{in},\alpha_{out}}^{out\leftarrow in}(\mathbf{r}_{bx},\mathbf{r}_{x})V_{bx}(r_{bx})\langle r_{a}J_{A}M_{A}|\chi_{a}^{(+)}m_{j_{A}}\rangle\langle r_{bx}J_{a}M_{a}|\phi_{a}M_{a}\rangle$$

$$\times \frac{\delta[\mathbf{r}_{b}-\mathbf{f}(\mathbf{r}_{bx},\mathbf{r}_{x})]}{r_{b}^{2}}\frac{\delta[\mathbf{r}_{a}-\mathbf{g}(\mathbf{r}_{bx},\mathbf{r}_{x})]}{r_{a}^{2}}d\Omega_{r_{bx}}d\Omega_{r_{x}},$$
(A7)

where

$$\langle \chi_b^{(-)} m_{j_b} | r_b J_b M_b \rangle = \frac{4\pi}{k_b r_b} i^{-\lambda_b} e^{i\sigma_{\lambda_b}} f_{\lambda_b}(k_b r_b) \sum_{m_{\lambda_b}} \langle \lambda_b m_{\lambda_b} j_b m_{j_b} | J_b M_b \rangle Y_{\lambda_b}^{m_{\lambda_b}}(\hat{k}_b), \tag{A8}$$

$$\langle r_a J_A M_A | \chi_a^{(+)} m_{j_A} \rangle = \frac{4\pi}{k_a r_a} i^{\lambda_a} e^{i\sigma_{\lambda_a}} f_{\lambda_a}(k_a r_a) \sum_{m_{\lambda_a}} \langle \lambda_a m_{\lambda_a} j_A m_{j_A} | J_A M_A \rangle Y_{\lambda_a}^{m_{\lambda_a}*}(\hat{k}_a), \tag{A9}$$

$$\langle r_{bx}J_aM_a|\phi_aM_a\rangle = \frac{u_{l_a}(r_{bx})}{r_{bx}},\tag{A10}$$

and

$$\mathcal{G}_{\alpha_{in},\alpha_{out}}^{\text{out}\leftarrow\text{in}}(\mathbf{r}_{bx},\mathbf{r}_{x}) = \sum_{LS} (2S+1)\sqrt{(2J_{a}+1)(2J_{A}+1)(2J_{x}+1)(2J_{b}+1)}} \begin{cases} l_{x} & s_{xA} & J_{x} \\ \lambda_{b} & j_{b} & J_{b} \\ L & S & J \end{cases} \begin{cases} l_{a} & s_{bx} & J_{a} \\ \lambda_{a} & j_{A} & J_{A} \\ L & S & J \end{cases} \\ \times (-)^{s_{bx}+2j_{A}+j_{x}+j_{b}}\sqrt{(2s_{xA}+1)(2s_{bx}+1)} \begin{cases} j_{A} & j_{x} & s_{xA} \\ j_{b} & S & s_{bx} \end{cases} \sum_{M=-L}^{L} \sum_{m_{i_{x}}m_{\lambda_{b}}} \sum_{m_{i_{a}}m_{\lambda_{a}}} \\ \times \langle l_{x}m_{l_{x}}'\lambda_{b}m_{\lambda_{b}}' | LM \rangle \langle l_{a}m_{l_{a}}'\lambda_{a}m_{\lambda_{a}}' | LM \rangle Y_{l_{x}}^{m_{l_{x}}'*}(\hat{r}_{x})Y_{\lambda_{b}}^{m_{\lambda_{b}}'*}(\hat{f})Y_{l_{a}}^{m_{l_{a}}'}(\hat{r}_{bx})Y_{\lambda_{a}}^{m_{\lambda_{a}}'}(\hat{g}). \end{cases}$$
(A11)

Here, I assume \hat{r}_{bx} is aligned with the *z* direction, so $\int d\Omega_{r_{bx}} = 4\pi$. In the zero-range approximation, I set $\hat{r}_x = \hat{r}_a = \hat{r}_b$, $r_b = cr_x$, and $r_a = r_x$, where *c* is the mass ratio of *A* to *B*^{*}. Also, I use the known relations

$$Y_{l_a}^{m'_{l_a}}(\hat{r}_{bx}) = \left(\frac{2l_a+1}{4\pi}\right)^{1/2} \delta_{m'_{l_a},0},\tag{A12}$$

$$Y_l^{m*}(\theta,\varphi) = (-)^m Y_{l,-m}(\theta,\varphi), \tag{A13}$$

$$\int Y_{l_x}^{m'_{l_x}*}(\hat{r}_x)Y_{\lambda_b}^{m'_{\lambda_b}*}(\hat{r}_x)Y_{\lambda_a}^{m'_{\lambda_a}}(\hat{r}_x) d\Omega_{r_x} = (-)^{l_x+\lambda_b-\lambda_a} \left[\frac{(2l_x+1)(2\lambda_b+1)}{4\pi(2\lambda_a+1)}\right]^{1/2} \langle l_x 0\lambda_b 0|\lambda_a 0\rangle \langle l_x m'_{l_x}\lambda_b m'_{\lambda_b}|\lambda_a m'_{\lambda_a}\rangle.$$
(A14)

Thus, the source term can be rewritten as

$$\langle r_{x}\beta M_{x}m_{j_{b}}|\rho M_{a}m_{j_{A}}\rangle = \frac{16\pi^{2}}{k_{a}k_{b}}\sum_{\alpha_{\text{out}},\alpha_{\text{in}}}\sum_{M_{J}}\sum_{M_{A}}\sum_{M_{b}}\sum_{m_{\lambda_{a}}}\sum_{m_{\lambda_{a}}}\delta_{\beta,\alpha_{\text{out}}}\langle J_{x}M_{x}J_{b}M_{b}|JM_{J}\rangle\langle J_{a}M_{a}J_{A}M_{A}|JM_{J}\rangle$$

$$\times \langle \lambda_{b}m_{\lambda_{b}}j_{b}m_{j_{b}}|J_{b}M_{b}\rangle\langle \lambda_{a}m_{\lambda_{a}}j_{A}m_{j_{A}}|J_{A}M_{A}\rangle Y_{\lambda_{a}}^{m_{\lambda_{a}}*}(\hat{k}_{a})Y_{\lambda_{b}}^{m_{\lambda_{b}}}(\hat{k}_{b})\Lambda(r_{x},\alpha_{\text{in}},\alpha_{\text{out}}), \tag{A15}$$

where

$$\Lambda(r_x, \alpha_{\rm in}, \alpha_{\rm out}) = i^{\lambda_a - \lambda_b} e^{i(\sigma_{\lambda_a} + \sigma_{\lambda_b})} \frac{f_{\lambda_b}(k_b c r_x)}{c r_x} \frac{f_{\lambda_a}(k_a r_x)}{r_x} \mathfrak{G}_{\alpha_{\rm in}, \alpha_{\rm out}} D_0, \tag{A16}$$

with

$$D_0 = (4\pi)^{1/2} \int r_{bx}^2 V_{bx}(r_{bx}) \langle r_{bx} J_a M_a | \phi_a M_a \rangle \, dr_{bx}, \tag{A17}$$

014609-9

$$\mathfrak{G}_{\alpha_{in},\alpha_{out}} = \left[\frac{(2l_x+1)(2\lambda_b+1)(2l_a+1)(2J_a+1)(2J_A+1)(2J_x+1)(2J_b+1)}{4\pi(2\lambda_a+1)}\right]^{1/2} (-1)^{l_x+\lambda_b-\lambda_a} \langle l_x 0\lambda_b 0 | \lambda_a 0 \rangle$$

$$\times \sum_{LS} (2S+1) \begin{cases} l_x & s_{xA} & J_x \\ \lambda_b & j_b & J_b \\ L & S & J \end{cases} \begin{cases} l_a & s_{bx} & J_a \\ \lambda_a & j_A & J_A \\ L & S & J \end{cases}$$

$$\times (-)^{s_{bx}+2j_A+j_x+j_b} \sqrt{(2s_{xA}+1)(2s_{bx}+1)} \begin{cases} j_A & j_x & s_{xA} \\ j_b & S & s_{bx} \end{cases} \sum_{m'_{\lambda_a}} \langle l_a 0\lambda_a m'_{\lambda_a} | Lm'_{\lambda_a} \rangle \delta_{\lambda_a,L}.$$
(A18)

It should be noted that in the zero-range approximation, the condition $\delta_{\lambda_a,L}$ is required. Thus, only the $l_a = 0$ zero-range case provides a good approximation to the finite-range results. For cases where $l_a \neq 0$, the zero-range source term yields zero, even when the finite-range source term is significant.

Then, the φ_x function in the partial wave basis with the zero-range approximation becomes

$$\langle r_{x}\beta M_{x}m_{j_{b}}|\varphi_{x}(\mathbf{k}_{b})M_{a}m_{j_{A}}\rangle = \sum_{\alpha_{\text{out}},\alpha_{\text{in}}}\sum_{M_{J}}\sum_{M_{A}}\sum_{M_{b}}\sum_{M_{\lambda_{b}}}\sum_{m_{\lambda_{a}}}\delta_{\beta,\alpha_{\text{out}}}\langle J_{x}M_{x}J_{b}M_{b}|JM_{J}\rangle\langle J_{a}M_{a}J_{A}M_{A}|JM_{J}\rangle$$

$$\times \langle \lambda_{b}m_{\lambda_{b}}j_{b}m_{j_{b}}|J_{b}M_{b}\rangle\langle \lambda_{a}m_{\lambda_{a}}j_{A}m_{j_{A}}|J_{A}M_{A}\rangle Y_{\lambda_{a}}^{m_{\lambda_{a}}*}(\hat{k}_{a})Y_{\lambda_{b}}^{m_{\lambda_{b}}}(\hat{k}_{b})\mathcal{R}(r_{x},\alpha_{\text{in}},\alpha_{\text{out}}),$$
(A19)

where

$$\mathcal{R}(r_x, \alpha_{\rm in}, \alpha_{\rm out}) = -\frac{32\pi^2 \mu_x}{\hbar^2 k_a k_b k_x} \frac{1}{c r_x} i^{\lambda_a - \lambda_b} e^{i(\sigma_{\lambda_a} + \sigma_{\lambda_b})} \mathfrak{G}_{\alpha_{\rm in}, \alpha_{\rm out}} D_0 \Pi(r_x), \tag{A20}$$

with

$$\Pi(r_x) = \int \frac{f_\beta(r'_x <) h_\beta^{(+)}(r'_x >)}{r'_x} f_{\lambda_b}(k_b c r'_x) f_{\lambda_a}(k_a r'_x) dr'_x.$$
(A21)

In the Vincent-Fortune method, under the zero-range approximation, I only consider the external part. For a given r_x ,

$$\Pi(r_x) = \Pi^{\text{int}}(r_x) + \Pi^{\text{ext}}(r_x).$$
(A22)

In the current approach, I use the finite-range method to compute Π^{int} and the zero-range method to compute Π^{ext} . For the asymptotic behavior, one has

$$f_{\lambda_b}(k_b c r_x) = \frac{i}{2} \Big[H_{\lambda_b}^{(-)}(k_b c r_x) - S_{\lambda_b} H_{\lambda_b}^{(+)}(k_b c r_x) \Big],$$
(A23)

$$f_{\lambda_a}(k_a r_x) = \frac{i}{2} \Big[H_{\lambda_a}^{(-)}(k_a r_x) - S_{\lambda_a} H_{\lambda_a}^{(+)}(k_a r_x) \Big], \tag{A24}$$

$$\Pi^{\text{ext}}(r_{x}) = -\frac{1}{4} \int_{R}^{\infty} \frac{dr'_{x}}{r'_{x}} f_{\beta}(r_{x}) \Big[H^{(+)}_{l_{x}}(k_{x}r'_{x})H^{(-)}_{\lambda_{b}}(k_{b}cr'_{x})H^{(-)}_{\lambda_{a}}(k_{a}r'_{x}) - S_{\lambda_{b}}H^{(+)}_{l_{x}}(k_{x}r'_{x})H^{(+)}_{\lambda_{b}}(k_{b}cr'_{x})H^{(-)}_{\lambda_{a}}(k_{a}r'_{x}) - S_{\lambda_{a}}H^{(+)}_{l_{x}}(k_{x}r'_{x})H^{(-)}_{\lambda_{b}}(k_{b}cr'_{x})H^{(-)}_{\lambda_{a}}(k_{a}r'_{x}) - S_{\lambda_{a}}H^{(+)}_{l_{x}}(k_{x}r'_{x})H^{(-)}_{\lambda_{b}}(k_{b}cr'_{x})H^{(-)}_{\lambda_{a}}(k_{a}r'_{x}) - S_{\lambda_{a}}H^{(+)}_{l_{x}}(k_{x}r'_{x})H^{(-)}_{\lambda_{b}}(k_{b}cr'_{x})H^{(-)}_{\lambda_{b}}(k_{b}cr'_{x})H^{(-)}_{\lambda_{a}}(k_{a}r'_{x}) - S_{\lambda_{a}}H^{(+)}_{l_{x}}(k_{x}r'_{x})H^{(-)}_{\lambda_{b}}(k_{b}cr'_{x})H^{(-)}_{\lambda_{a}}(k_{a}r'_{x}) - S_{\lambda_{a}}H^{(+)}_{l_{x}}(k_{x}r'_{x})H^{(+)}_{\lambda_{b}}(k_{b}cr'_{x})H^{(-)}_{\lambda_{a}}(k_{a}r'_{x}) - S_{\lambda_{a}}H^{(+)}_{l_{x}}(k_{x}r'_{x})H^{(-)}_{\lambda_{b}}(k_{b}cr'_{x})H^{(-)}_{\lambda_{a}}(k_{a}r'_{x}) - S_{\lambda_{a}}H^{(+)}_{l_{x}}(k_{x}r'_{x})H^{(+)}_{\lambda_{b}}(k_{b}cr'_{x})H^{(-)}_{\lambda_{a}}(k_{a}r'_{x}) - S_{\lambda_{a}}H^{(+)}_{\lambda_{b}}(k_{b}cr'_{x})H^{(+)}_{\lambda_{b}}(k_{b}cr'_{x})H^{(-)}_{\lambda_{a}}(k_{a}r'_{x}) - S_{\lambda_{a}}H^{(+)}_{\lambda_{b}}(k_{b}cr'_{x})H^{(+)}_{\lambda_{b}}(k_{b}cr'_{x})H^{(+)}_{\lambda_{a}}(k_{a}r'_{x}) - S_{\lambda_{a}}H^{(+)}_{\lambda_{b}}(k_{b}cr'_{x})H^{(+)}_{\lambda_{b}}(k_{b}cr'_{x})H^{(+)}_{\lambda_{a}}(k_{a}r'_{x}) - S_{\lambda_{b}}H^{(+)}_{\lambda_{b}}(k_{b}cr'_{x})H^{(+)}_{\lambda_{b}}(k_{b}cr'_{x})H^{(+)}_{\lambda_{a}}(k_{a}r'_{x}) - S_{\lambda_{b}}H^{(+)}_{\lambda_{b}}(k_{b}cr'_{x})H^{(+)}_{\lambda_{b}}(k_{b}cr'_{x})H^{(+)}_{\lambda_{b}}(k_{b}cr'_{x}) - S_{\lambda_{b}}H^{(+)}_{\lambda_{b}}(k_{b}cr'_{x})H^{(+)}_{\lambda_{b}}(k_{b}cr'_{x})H^{(+)}_{\lambda_{b}}(k_{b}cr'_{x}) - S_{\lambda_{b}}H^{(+)}_{\lambda_{b}}(k_{b}cr'_{x})H^{(+)}_{\lambda_{b}}(k_{b}cr'_{x})H^{(+)}_{\lambda_{b}}(k_{b}cr'_{x}) - S_{\lambda_{b}}H^{(+)}_{\lambda_{b}}(k_{b}cr'_{x})H^{(+)}_{\lambda_{b}}(k_{b}cr'_{x}) - S_{\lambda_{b}}H^{(+)}_{\lambda_{b}}(k_{b}cr'_{x})H^{(+)}_{\lambda_{b}}(k_{b}cr'_{x}) - S_{\lambda_{b}}H^{(+)}_{\lambda_{b}}(k_{b}cr'_{x}) - S_{\lambda_{b}}H^{(+)}_{\lambda_{b}}(k_{b}cr'_{x}) - S_{\lambda_{b}}H^{(+)}_{\lambda_{b}}(k_{b}cr'_{x}) - S_{\lambda_{b}}H^{(+)}_{\lambda_{b}}(k_{b}cr'_{x}) - S_{\lambda_{b}}H^{(+)}_{\lambda_{b}}(k_{b}cr'_{x}) - S_{\lambda_{b}}H^{(+)}_{\lambda_{b}}(k_{b}cr'_{x}) - S_$$

I treat the above equation in the complex plane to overcome the divergence of the integration on the real axis. For example, I consider the first product term:

$$H_{l_x}^{(+)}(k_x z)H_{\lambda_b}^{(-)}(k_b cz)H_{\lambda_a}^{(-)}(k_a z) \xrightarrow{z \to \infty} e^{ik_x z} e^{-ik_b cz} e^{-ik_a z} = e^{i(k_x - ck_b - k_a)(x + iy)} = e^{i(k_x - ck_b - k_a)x} e^{-(k_x - ck_b - k_a)y}.$$
(A26)

Now, if $-(k_x - ck_b - k_a) > 0$, the second exponential will diverge unless y takes a negative value. Thus, in the complex analysis, one should evaluate this term in the lower half of the complex plane.

Proceeding in a similar manner with the other terms, one has two groups of product terms, L(z) and U(z), where one is analytic in the lower half and the other in the upper half of the complex plane:

$$L(z) = \frac{1}{z} \Big[H_{l_x}^{(+)}(k_x z) H_{\lambda_b}^{(-)}(k_b c z) H_{\lambda_a}^{(-)}(k_a z) - S_{\lambda_b} H_{l_x}^{(+)}(k_x z) H_{\lambda_b}^{(+)}(k_b c z) H_{\lambda_a}^{(-)}(k_a z) \Big],$$

$$U(z) = \frac{1}{z} \Big[-S_{\lambda_a} H_{l_x}^{(+)}(k_x z) H_{\lambda_b}^{(-)}(k_b c z) H_{\lambda_a}^{(+)}(k_a z) + S_{\lambda_a} S_{\lambda_b} H_{l_x}^{(+)}(k_x z) H_{\lambda_b}^{(+)}(k_b c z) H_{\lambda_a}^{(+)}(k_a z) \Big].$$
(A27)

Moreover, as $|z| \to \infty$, $L(z) \to 0$ and $U(z) \to 0$. I can define three integration paths: C_1 is the real axis, C_2 is in the upper plane, and C_3 is in the lower plane:

$$\Pi^{\text{ext}}(r_x) = -\frac{f_{\beta}(r_x)}{4} \int_{C_1} dz L(z) - \frac{f_{\beta}(r_x)}{4} \int_{C_1} dz U(z)$$

$$= -\frac{f_{\beta}(r_x)}{4} \int_{C_2} dz U(z) - \frac{f_{\beta}(r_x)}{4} \int_{C_2'} dz U(z) - \frac{f_{\beta}(r_x)}{4} \int_{C_3} dz L(z) - \frac{f_{\beta}(r_x)}{4} \int_{C_3'} dz L(z)$$

$$= -\frac{f_{\beta}(r_x)}{4} \int_{C_2} dz U(z) - \frac{f_{\beta}(r_x)}{4} \int_{C_3} dz L(z).$$
(A28)

Here, I used the fact that, if the radius of the semicircle is taken to be large, then $\int_{C'_2} dz U(z) = 0$ and $\int_{C'_3} dz L(z) = 0$. Now,

$$C_2 := R + iy, \quad 0 \le y < \infty,$$

$$C_3 := R - iy, \quad 0 \le y < \infty.$$
(A29)

Therefore, one has

$$\Pi^{\text{ext}}(r_x) = -\frac{f_{\beta}(r_x)}{4} \bigg[i \int_0^\infty U(R+iy) \, dy - i \int_0^\infty L(R-iy) \, dy \bigg], \tag{A30}$$

with $U(R_{\max} + iy) \rightarrow 0$ and $L(R_{\max} - iy) \rightarrow 0$ as $y \rightarrow \infty$.

- G. Baur, F. Rösel, D. Trautmann, and R. Shyam, Phys. Rep. 111, 333 (1984).
- [2] I. Tanihata, Prog. Part. Nucl. Phys. 35, 505 (1995).
- [3] G. F. Grinyer, D. Bazin, A. Gade, J. A. Tostevin, P. Adrich, M. D. Bowen, B. A. Brown, C. M. Campbell, J. M. Cook, T. Glasmacher, S. McDaniel, P. Navrátil, A. Obertelli, S. Quaglioni, K. Siwek, J. R. Terry, D. Weisshaar, and R. B. Wiringa, Phys. Rev. Lett. **106**, 162502 (2011).
- [4] T. Aumann, C. Barbieri, D. Bazin, C. Bertulani, A. Bonaccorso, W. Dickhoff, A. Gade, M. Gómez-Ramos, B. Kay, A. Moro, T. Nakamura, A. Obertelli, K. Ogata, S. Paschalis, and T. Uesaka, Prog. Part. Nucl. Phys. **118**, 103847 (2021).
- [5] Z. H. Yang, Y. Kubota, A. Corsi, K. Yoshida, X.-X. Sun, J. G. Li, M. Kimura, N. Michel, K. Ogata, C. X. Yuan, Q. Yuan, G. Authelet, H. Baba, C. Caesar, D. Calvet, A. Delbart, M. Dozono, J. Feng, F. Flavigny, J.-M. Gheller *et al.*, Phys. Rev. Lett. **126**, 082501 (2021).
- [6] P. J. Li, D. Beaumel, J. Lee, M. Assié, S. Chen, S. Franchoo, J. Gibelin, F. Hammache, T. Harada, Y. Kanada-En'yo, Y. Kubota, S. Leblond, P. F. Liang, T. Lokotko, M. Lyu, F. M. Marqués, Y. Matsuda, K. Ogata, H. Otsu, E. Rindel *et al.*, Phys. Rev. Lett. 131, 212501 (2023).
- [7] S. Pandit, A. Shrivastava, K. Mahata, N. Keeley, V. Parkar, R. Palit, P. Rout, K. Ramachandran, A. Kumar, S. Bhattacharyya, V. Nanal, S. Biswas, S. Saha, J. Sethi, P. Singh, and S. Kailas, Phys. Lett. B 820, 136570 (2021).
- [8] J. E. Escher, J. T. Harke, F. S. Dietrich, N. D. Scielzo, I. J. Thompson, and W. Younes, Rev. Mod. Phys. 84, 353 (2012).
- [9] A. Tumino, C. A. Bertulani, M. La Cognata, L. Lamia, R. G. Pizzone, S. Romano, and S. Typel, Annu. Rev. Nucl. Part. Sci. 71, 345 (2021).
- [10] A. Tumino, C. Bertulani, S. Cherubini, G. D'Agata, A. D. Pietro, P. Figuera, G. Guardo, M. Gulino, S. Hayakawa, M.

La Cognata, M. L. Commara, L. Lamia, D. Lattuada, M. Mazzocco, A. Moro, J. Mraźek, A. Oliva, S. Palmerini, R. Pizzone, G. Rapisarda *et al.*, Prog. Part. Nucl. Phys. **143**, 104164 (2025).

- [11] J. Lei, Phys. Rev. C 111, 034610 (2025).
- [12] L. Faddeev, Mathematical Aspects of the Three-body Problem in the Quantum Scattering Theory, Academy of Sciences of the U.S.S.R.: Works of the Steklov Mathematical Institute Vol. 69 (Israel Program for Scientific Translations, 1965)
- [13] I. Thompson, F. Nunes, and B. Danilin, Comput. Phys. Commun. 161, 87 (2004).
- [14] A. Deltuva, Phys. Rev. C 88, 011601(R) (2013).
- [15] A. M. Mukhamedzhanov, V. Eremenko, and A. I. Sattarov, Phys. Rev. C 86, 034001 (2012).
- [16] N. Austern, M. Yahiro, and M. Kawai, Phys. Rev. Lett. 63, 2649 (1989).
- [17] N. Austern, Y. Iseri, M. Kamimura, M. Kawai, G. Rawitscher, and M. Yahiro, Phys. Rep. 154, 125 (1987).
- [18] G. Baur and D. Trautmann, Phys. Rep. 25, 293 (1976).
- [19] R. Chatterjee, P. Banerjee, and R. Shyam, Nucl. Phys. A 692, 476 (2001).
- [20] M. Ichimura, N. Austern, and C. M. Vincent, Phys. Rev. C 32, 431 (1985).
- [21] T. Udagawa and T. Tamura, Phys. Rev. C 24, 1348 (1981).
- [22] M. Hussein and K. McVoy, Nucl. Phys. A 445, 124 (1985).
- [23] G. Potel, F. M. Nunes, and I. J. Thompson, Phys. Rev. C 92, 034611 (2015).
- [24] B. V. Carlson, R. Capote, and M. Sin, Few-Body Syst. 57, 307 (2016).
- [25] F. Torabi and B. V. Carlson, J. Phys. G 50, 045107 (2023).
- [26] J. Lei and A. M. Moro, Phys. Rev. C 92, 044616 (2015).
- [27] J. Lei and A. M. Moro, Phys. Rev. C 92, 061602(R) (2015).
- [28] J. Lei and A. M. Moro, Phys. Rev. C 95, 044605 (2017).

- [29] J. Lei and A. M. Moro, Phys. Rev. C 97, 011601(R) (2018).
- [30] G. Potel, G. Perdikakis, B. V. Carlson, M. C. Atkinson, W. H. Dickhoff, J. E. Escher, M. S. Hussein, J. Lei, W. Li, A. O. Macchiavelli, A. M. Moro, F. M. Nunes, S. D. Pain, and J. Rotureau, Eur. Phys. J. A 53, 178 (2017).
- [31] R. Huby and J. R. Mines, Rev. Mod. Phys. 37, 406 (1965).
- [32] I. J. Thompson, J. Phys.: Conf. Ser. 312, 082041 (2011).
- [33] C. M. Vincent and H. T. Fortune, Phys. Rev. C 2, 782 (1970).
- [34] J. Lei and A. M. Moro, Phys. Rev. Lett. 123, 232501 (2019).
- [35] J. Lei and A. M. Moro, Phys. Rev. C 108, 034612 (2023).
- [36] C. M. Vincent, Phys. Rev. 175, 1309 (1968).
- [37] J. Lei, Phys. Rev. C 97, 034628 (2018).
- [38] Y. Han, Y. Shi, and Q. Shen, Phys. Rev. C 74, 044615 (2006).
- [39] A. Koning and J. Delaroche, Nucl. Phys. A 713, 231 (2003).